

Investigating Flux Densities and Polarization of Methanol Masers in the Orion Nebula



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- ACCESS



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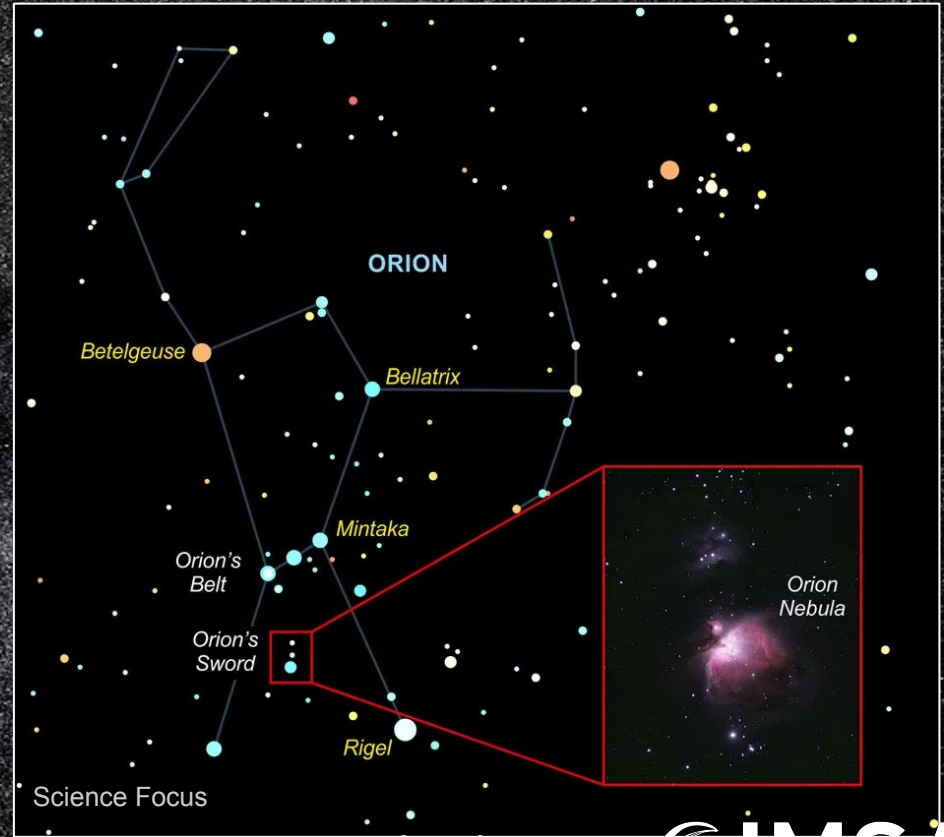
Background

- VLA (Very Large Array)
 - Radio Telescope
 - New Mexico, U.S.
- VOLS (VLA Orion-A Large Survey)
 - 20 Epochs
 - From April 3, 2022
 - To July 1, 2022



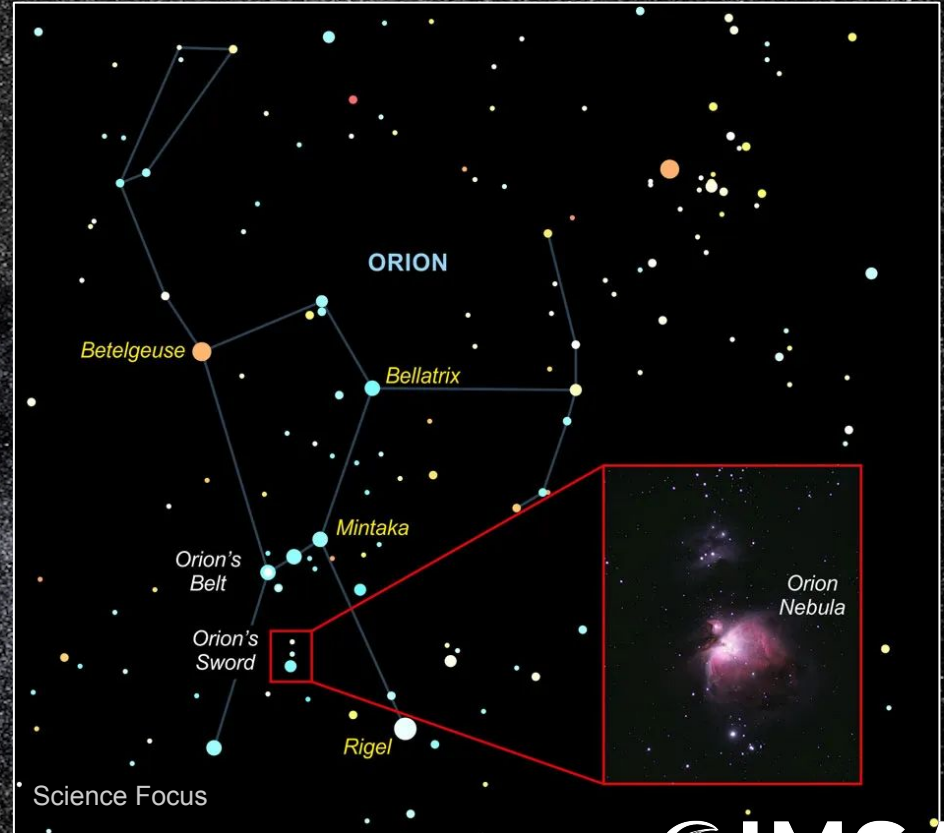
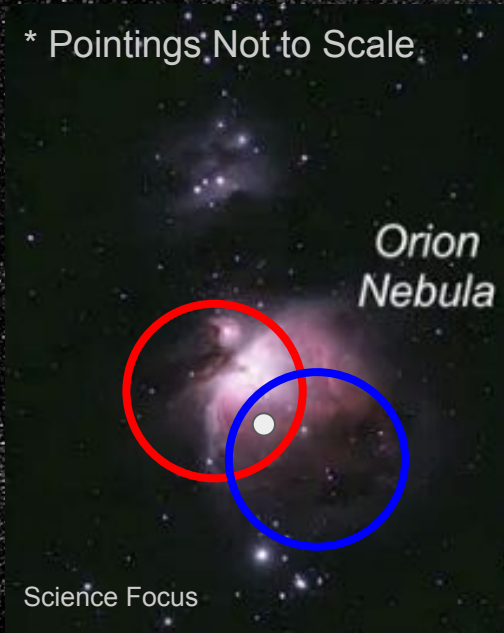
Background

■ Orion Nebula



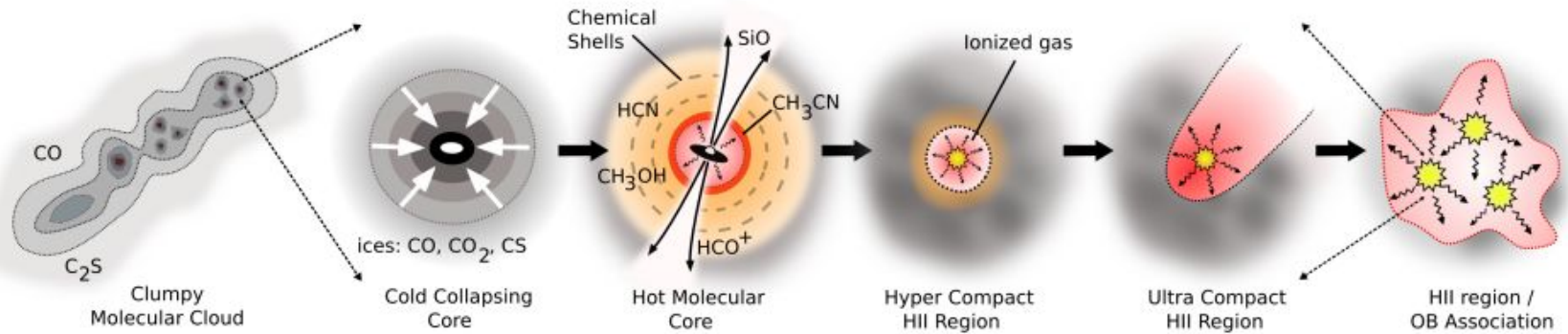
Background

- Orion Nebula
- Pointings



Background

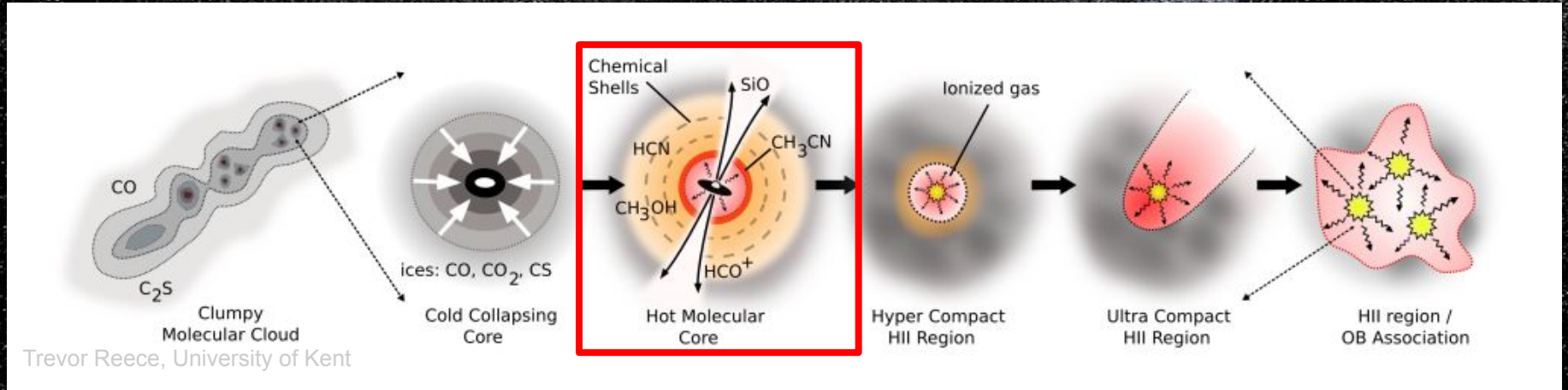
■ High Mass Star Formation



Trevor Reece, University of Kent

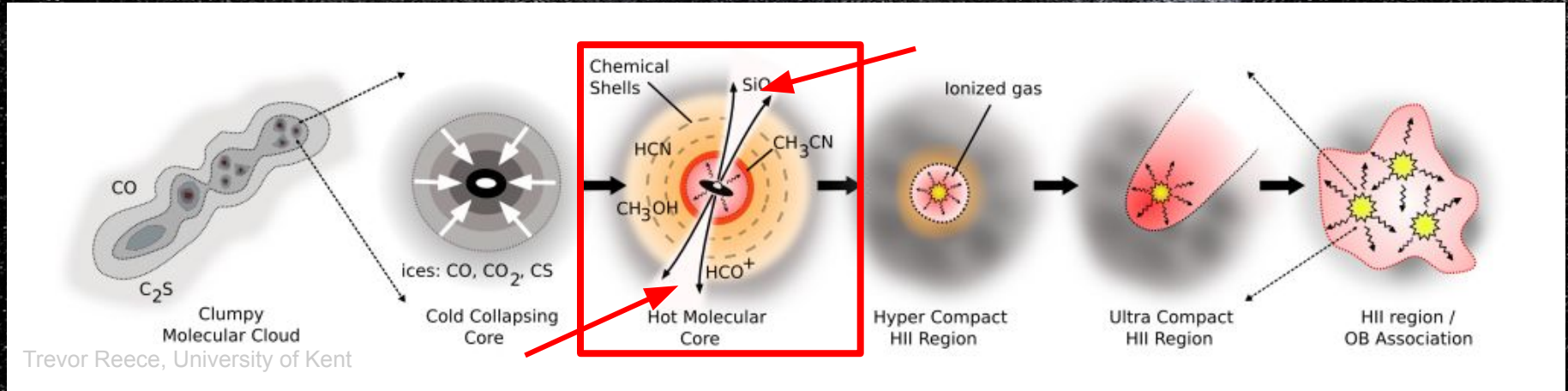
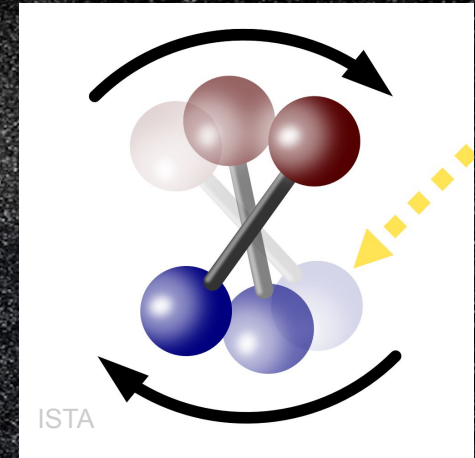
Background

- High Mass Star Formation
 - Molecular Outflow



Background

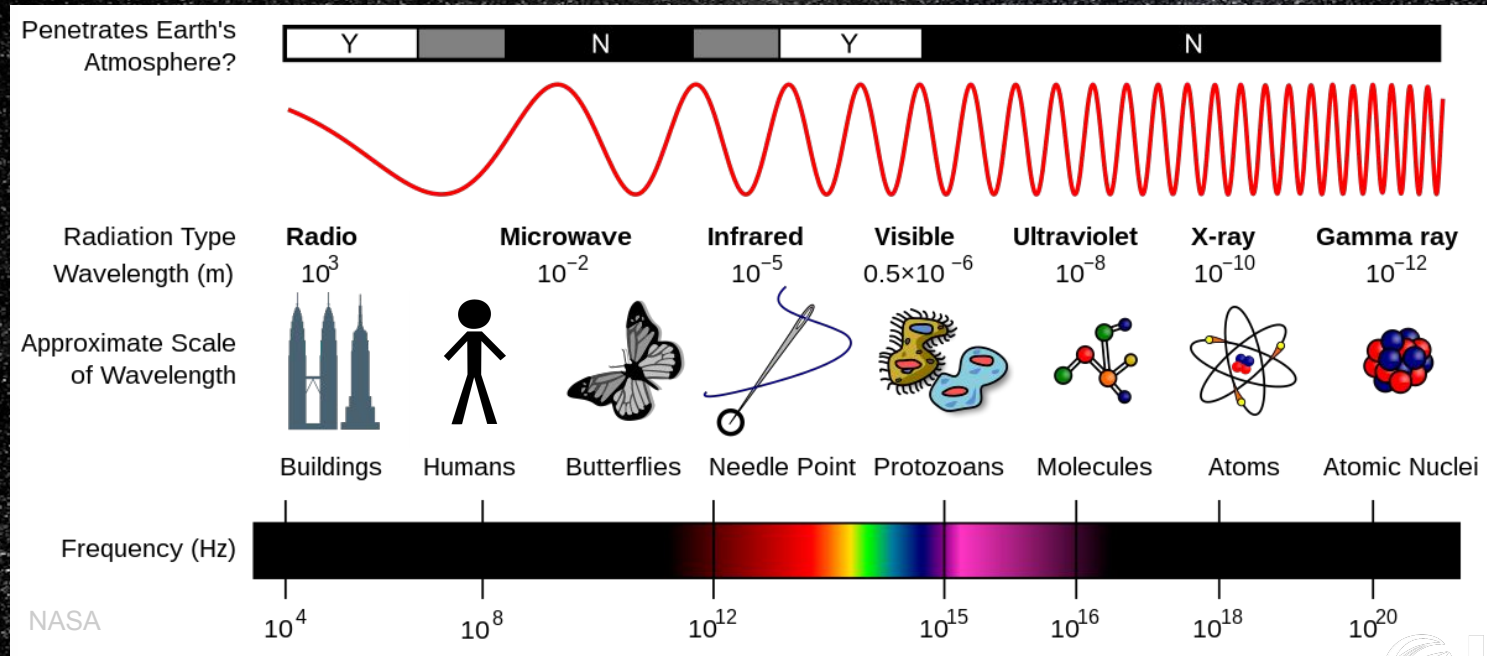
- High Mass Star Formation
 - Molecular Outflow
 - Masers



Background

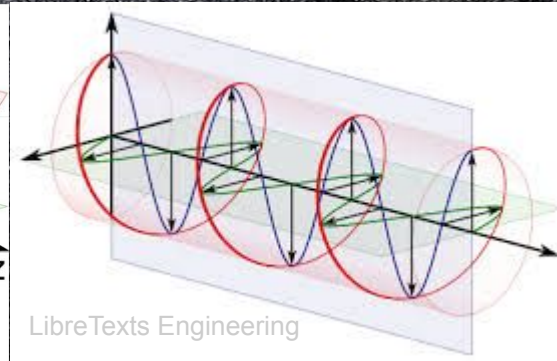
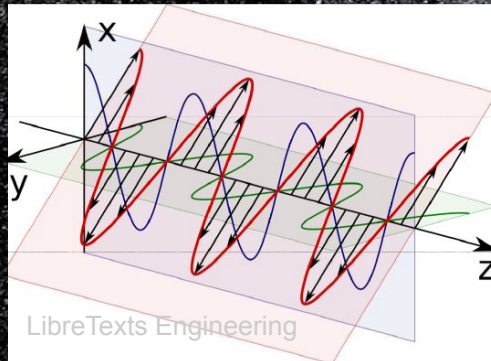
- Microwave Radiation
- Flux Density

$$1 \text{ Jy} = 10^{-26} \frac{\text{W}}{\text{m}^2 \text{ Hz}}$$



Background

- Light Wave Polarization
 - Stokes Intensity - IQUV
 - I - Overall Signal Intensity
 - Q, U - Linear Polarization, 45 degree increments
 - V - Circular Polarization, Left/Right-Hand



GeeksForGeeks

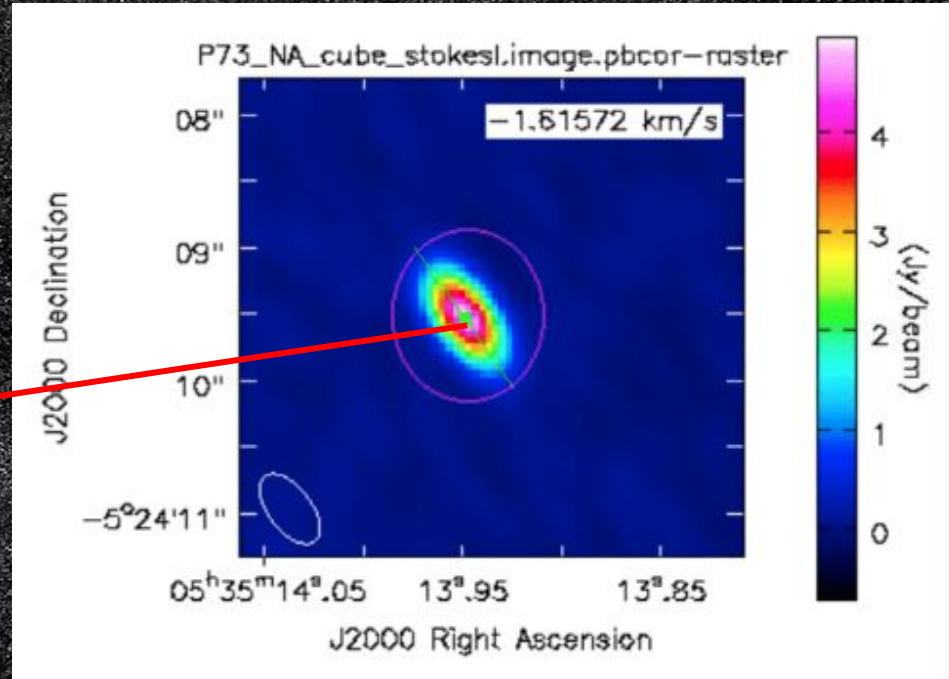
The diagram shows the Lewis structure of Methanol (CH₃OH). A central Carbon (C) atom is bonded to three Hydrogen (H) atoms and one Oxygen (O) atom. The Oxygen atom is further bonded to a Hydrogen (H) atom. All bonds are represented by single lines. The Carbon and Oxygen atoms are labeled with their respective symbols, and the Hydrogen atoms are labeled with 'H'. The structure is shown in a 3D perspective.

Methanol



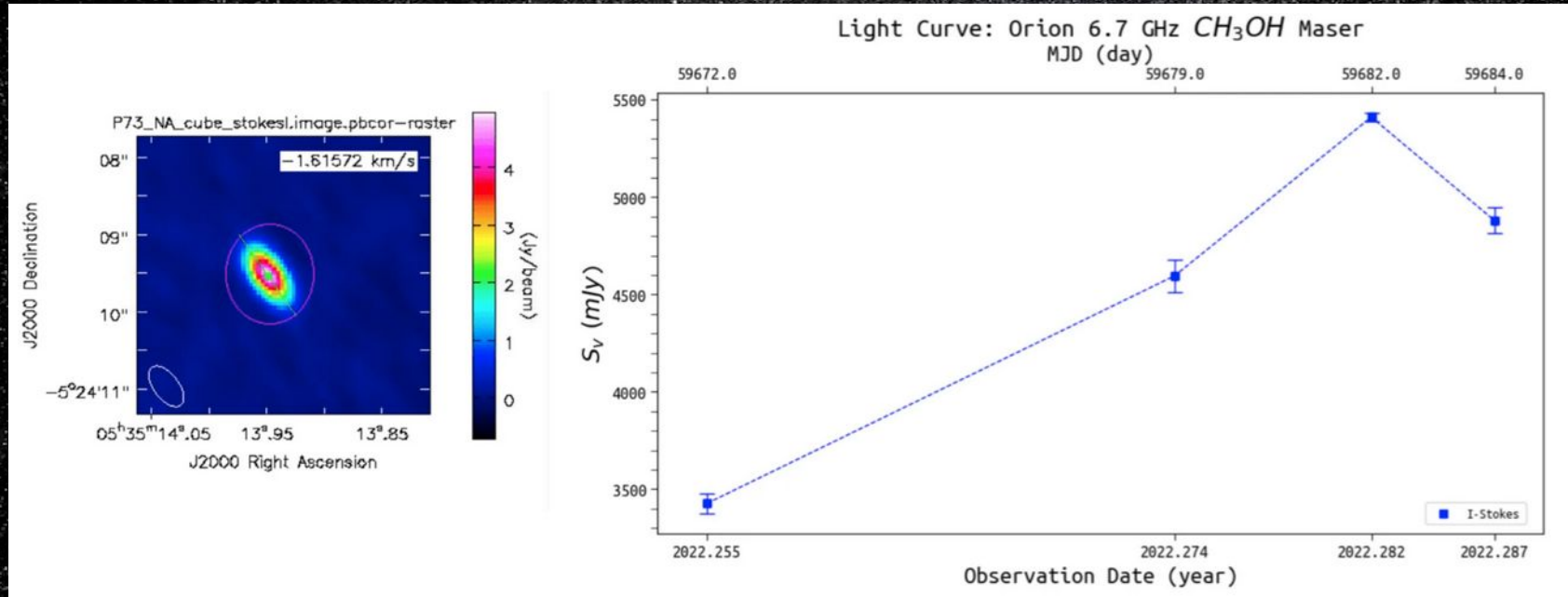
Background

■ Methanol (CH_3OH) Masers in Orion



Background

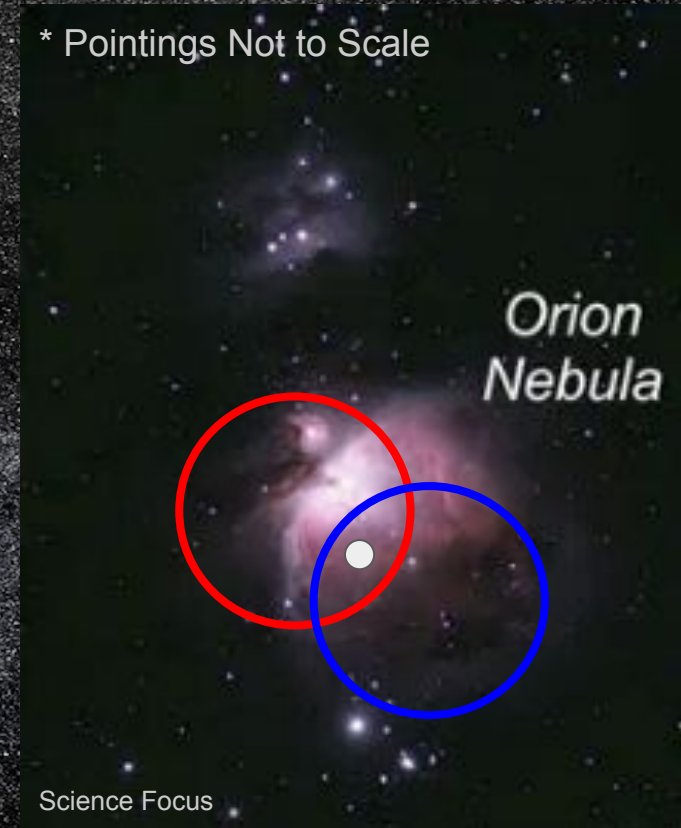
■ Methanol (CH_3OH) Masers in Orion



Sojka et al. (2024, AAS 244, 302.06)

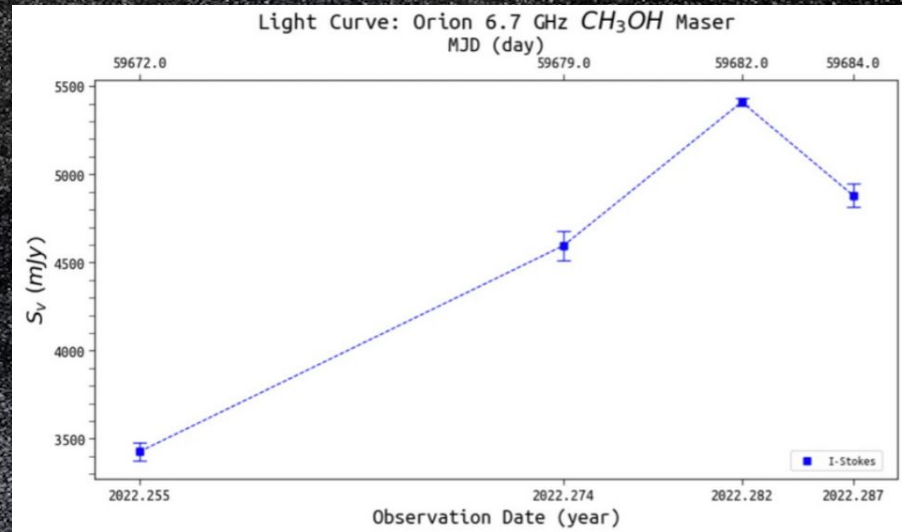
Objectives

- Obj 1: Confirmation using pointings in the night sky



Objectives

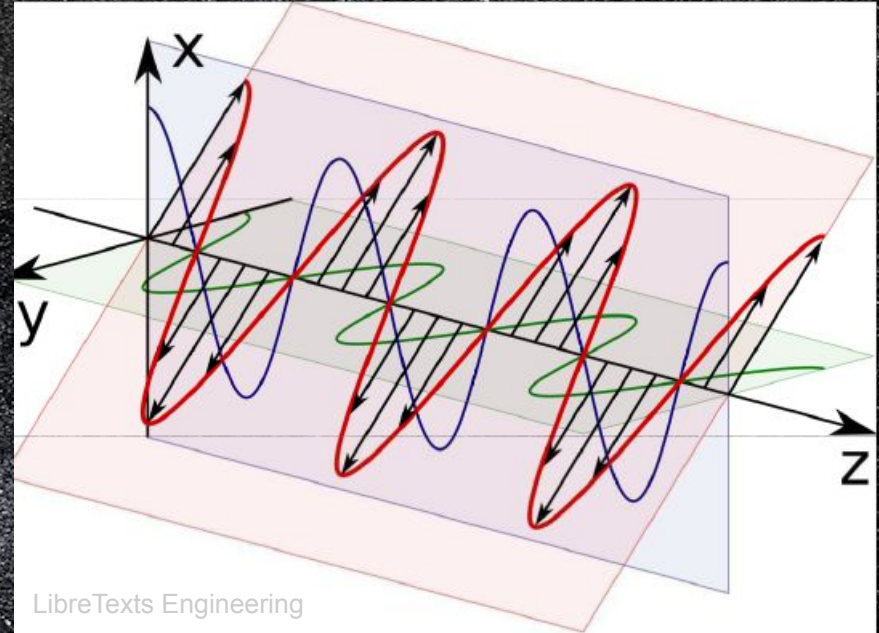
- Obj 1: Confirmation using pointings in the night sky
- Obj 2: Look for fast variability in the maser



Sojka et al. (2024, AAS 244, 302.06)

Objectives

- Obj 1: Confirmation using pointings in the night sky
- Obj 2: Look for fast variability in the maser
- Obj 3: Check for linear polarization (as well as circular polarization)



Methodology

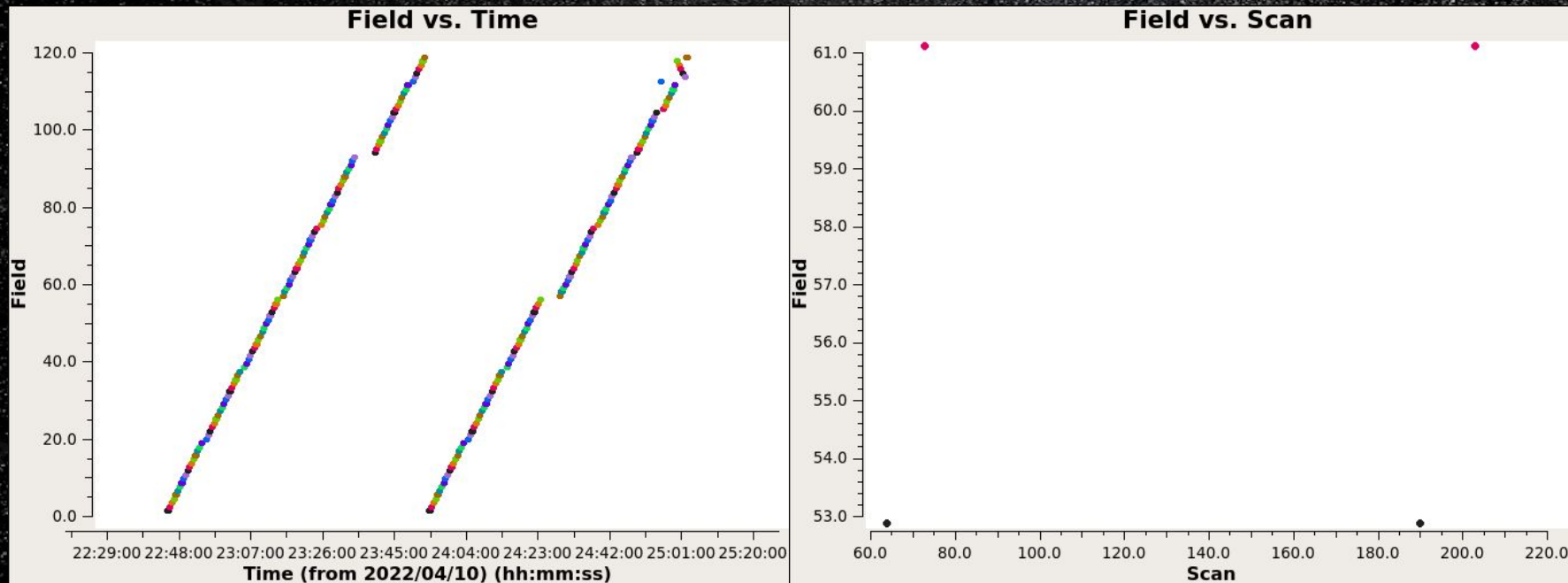
- NSF ACCESS, Jetstream2, University of Indiana
- CASA (Common Astronomical Software Applications)



Methodology

Epoch 2

■ Sky Scanning

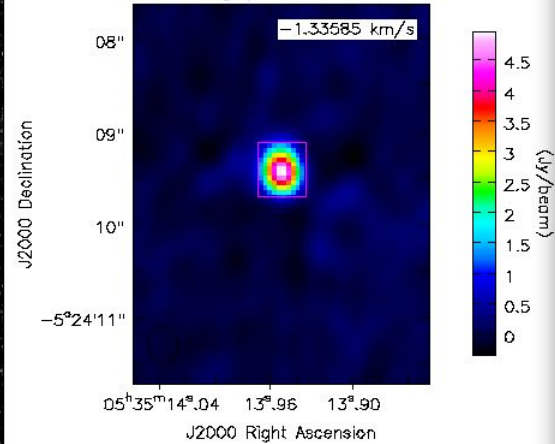


Methodology

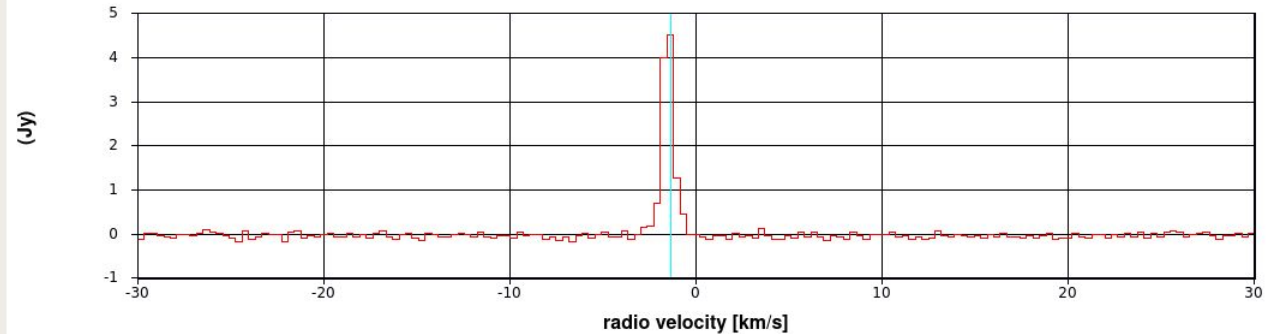
Epoch 2, Pointing 73

- Spectral Imaging
- Spectral Lines

cube_stokesI_Maser_A.image.pbcor--raster



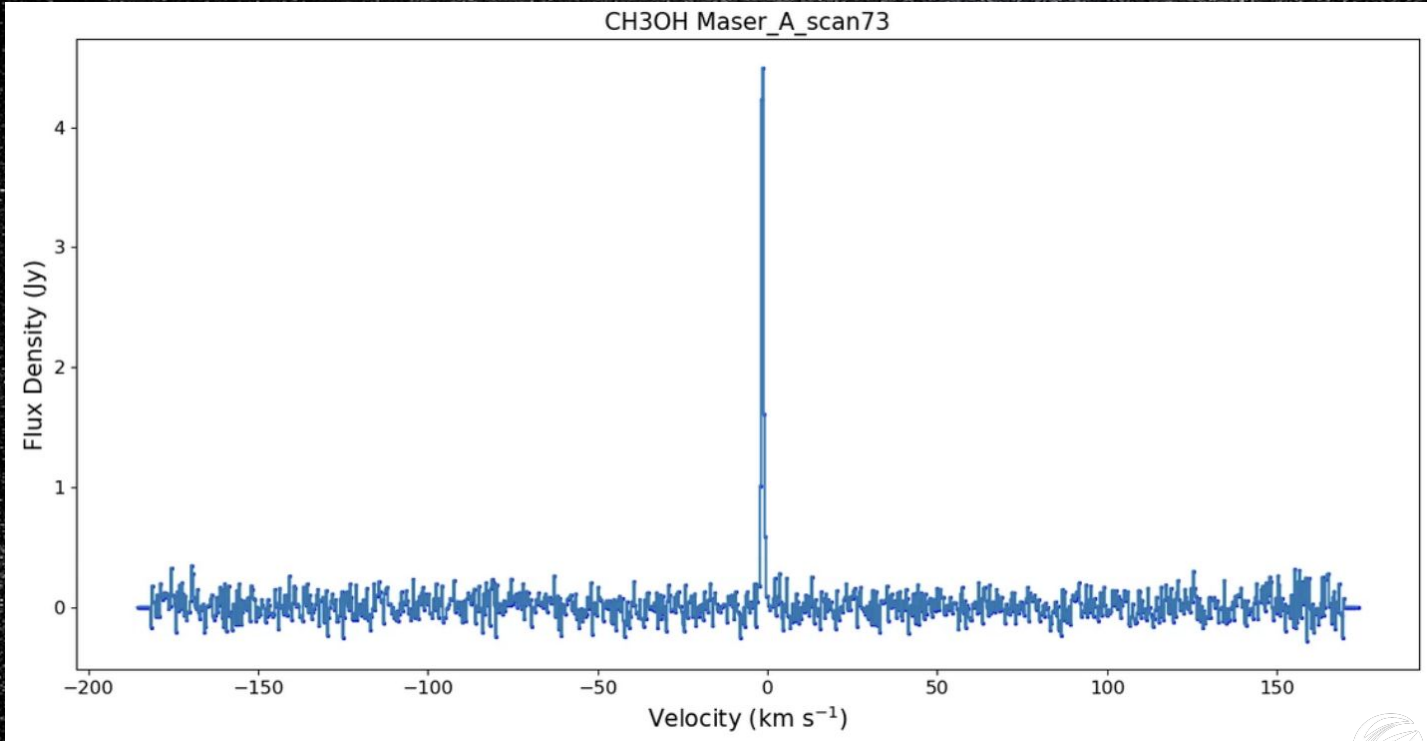
Rectangle Region Profile



Methodology

Epoch 2, Scan 73

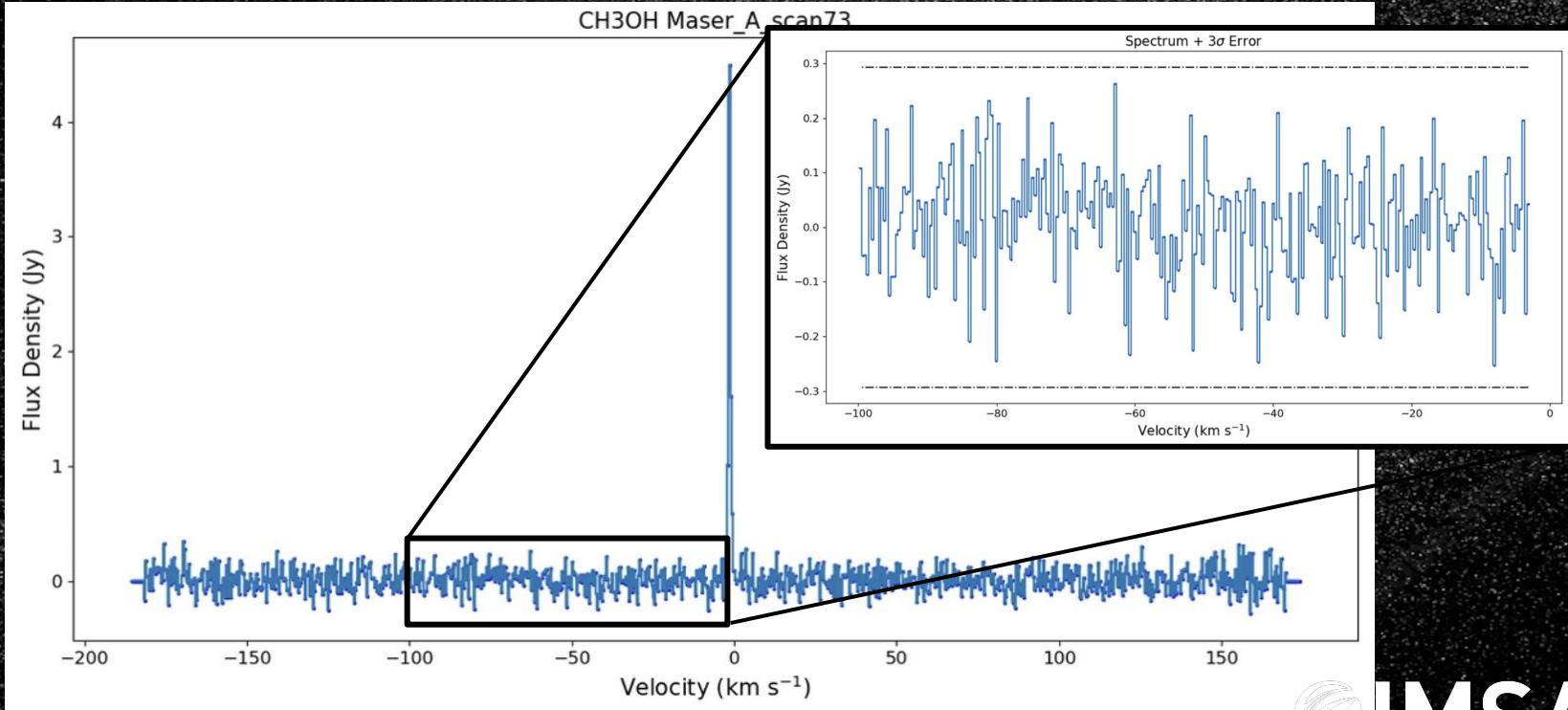
■ Examples of Signal and Noise Graphs



Methodology

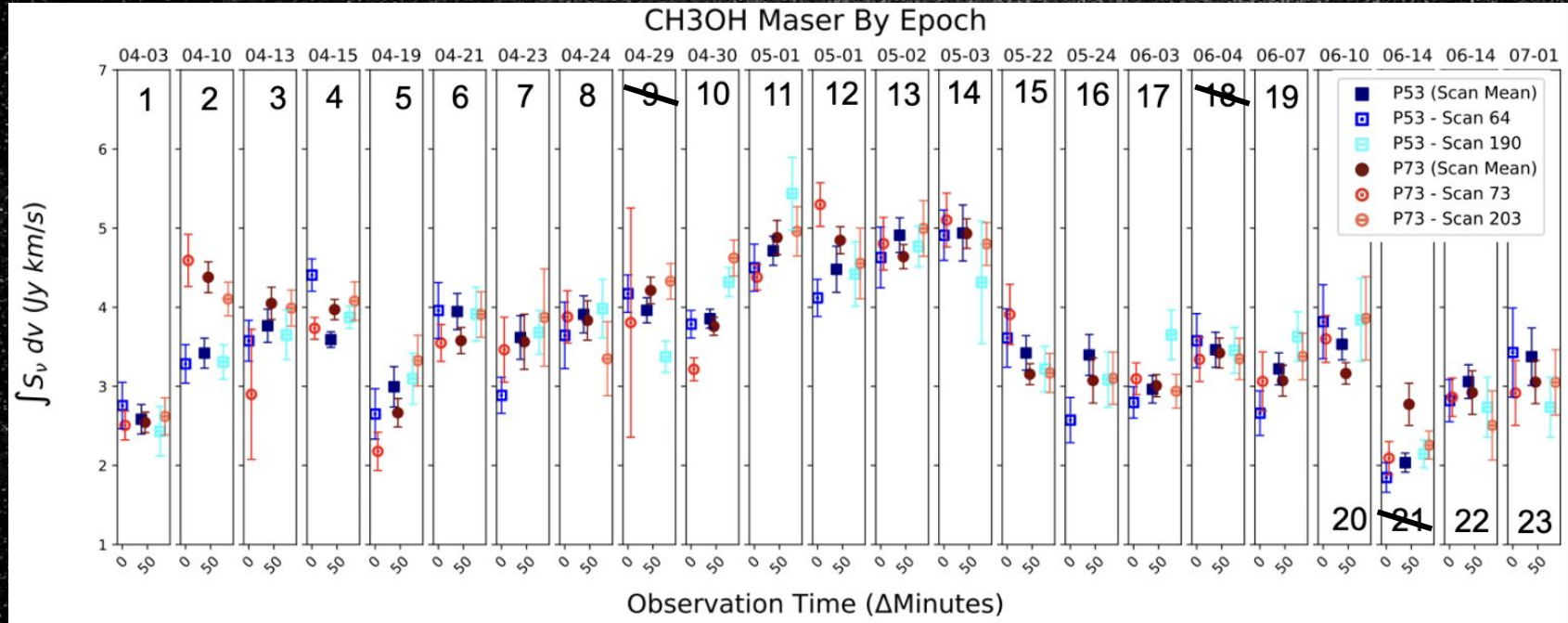
Epoch 2, Scan 73

■ Examples of Signal and Noise Graphs



Results and Discussion

■ Combined Light Curve of 23 epochs



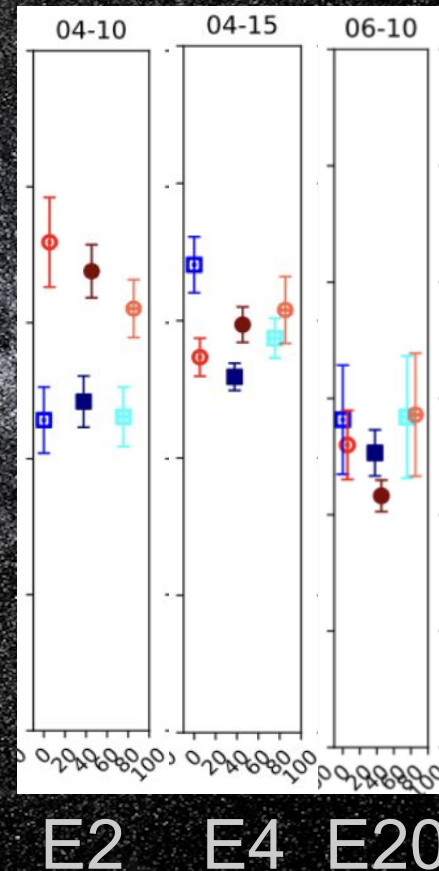
All Epochs

Objective 1

Epoch 2, Epoch 4

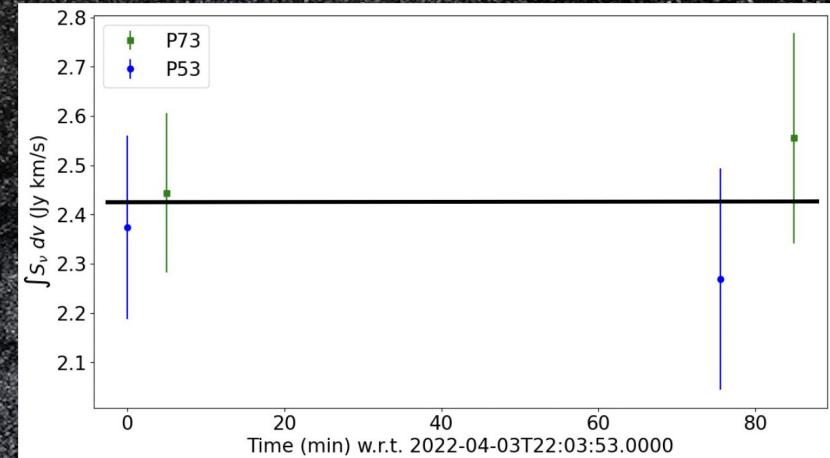
■ Overall Findings

- Flux densities consistent between different pointings for all epochs except epochs 2, 4, and 20
- General consistency for 17/20 epochs => general maser consistency



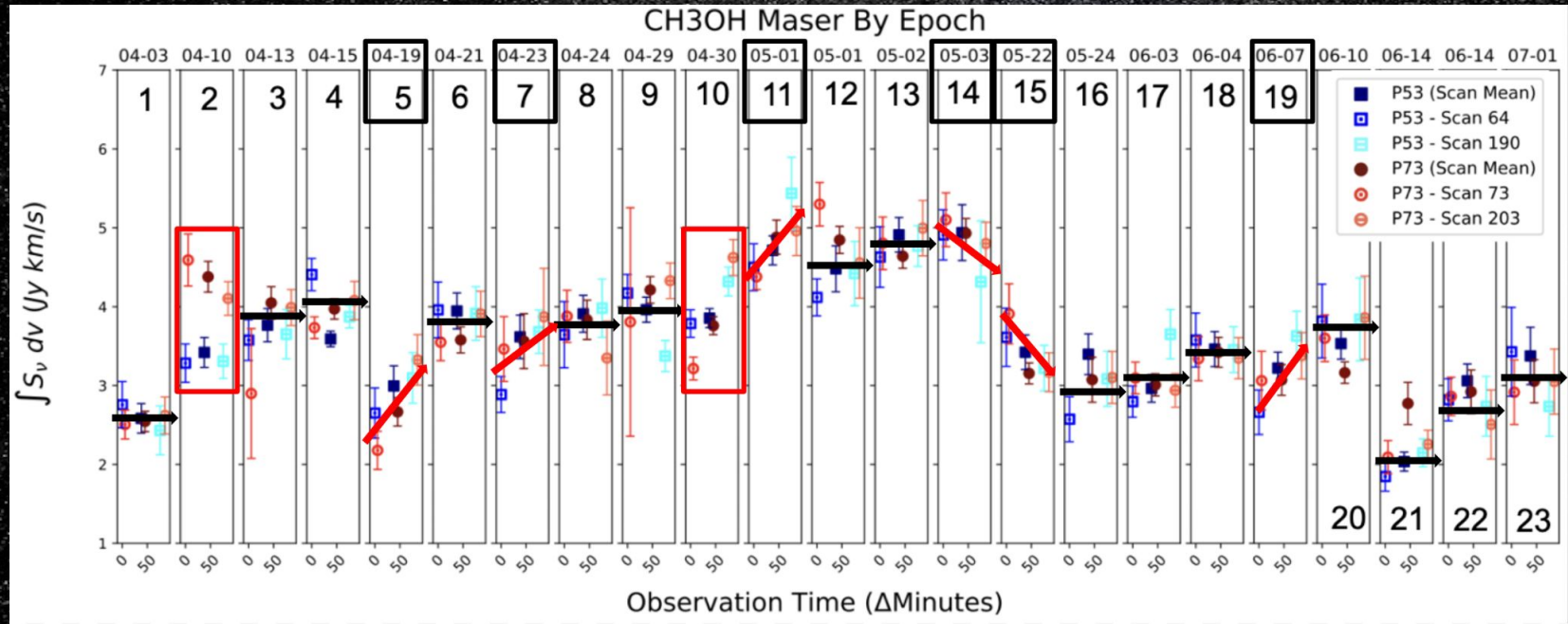
Objective 2

- Overall Findings
 - Possible evidence of fast variability in epochs with general trends
 - Generally similar scan values for most epochs
 - Offset scans in some epochs due to lack of telescope data or high noise



Objective 2

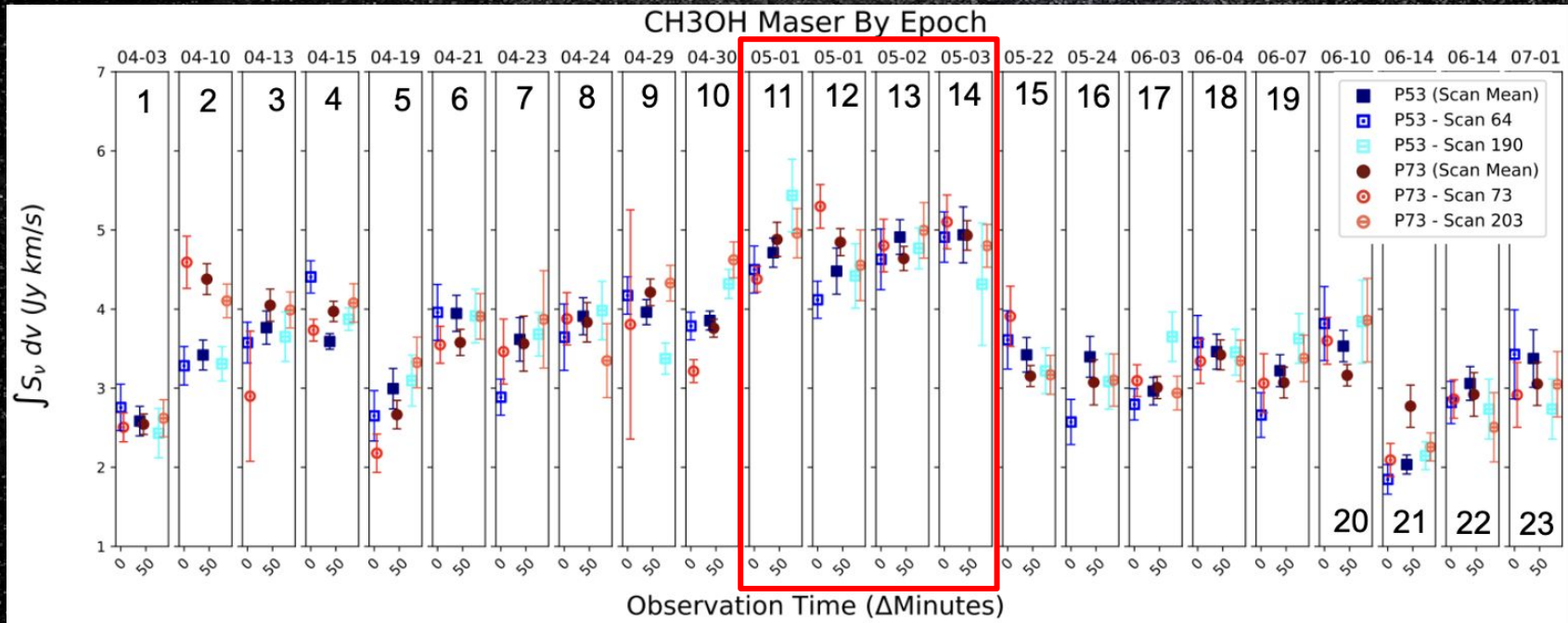
■ Fast Variability in the 20 Epochs



All Epochs

Discrepancy

■ May 1st to 3rd Flare

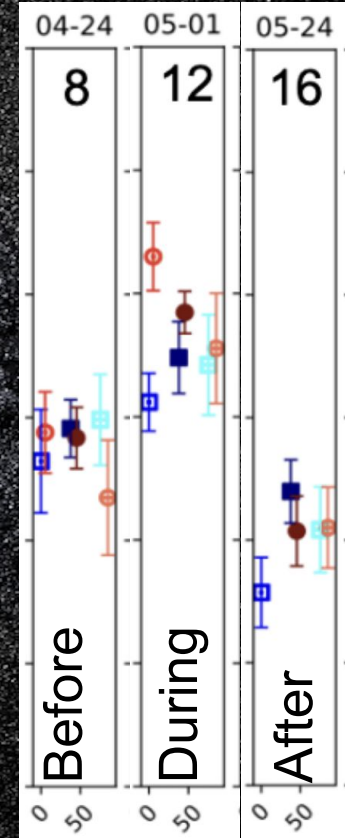


Epochs 11-14

Discrepancy

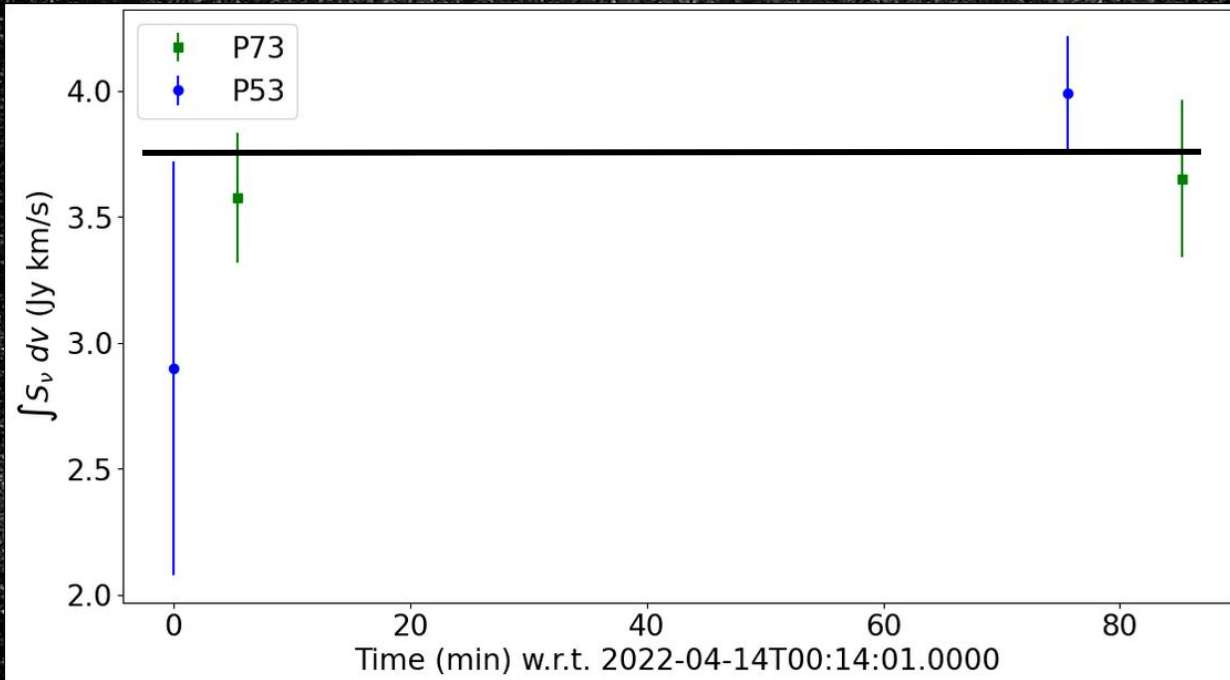
Epochs 8, 12, 16

- May 1st to 3rd Flare
 - Significant increase in flux density in epochs 11 through 14
 - Possible Causes:
 - Nearby young stellar object's accretion disk grows, leading to a larger radiation field affecting maser
 - Shockwave front passes through maser



■ Comparing Pointings

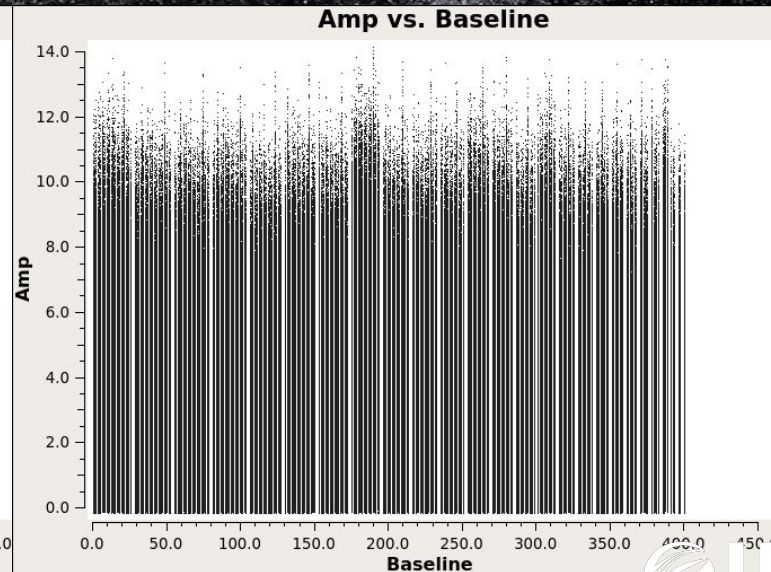
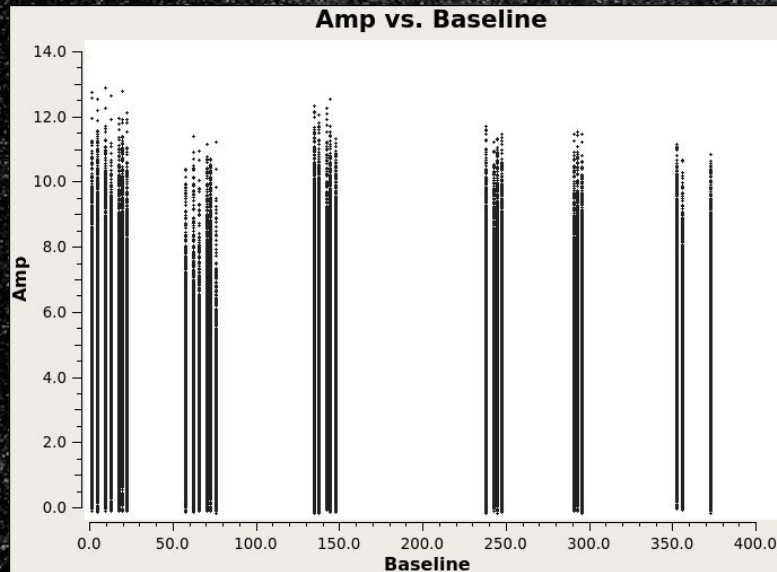
- Epoch 3 pointings are consistent, but one scan is offset



Discrepancy

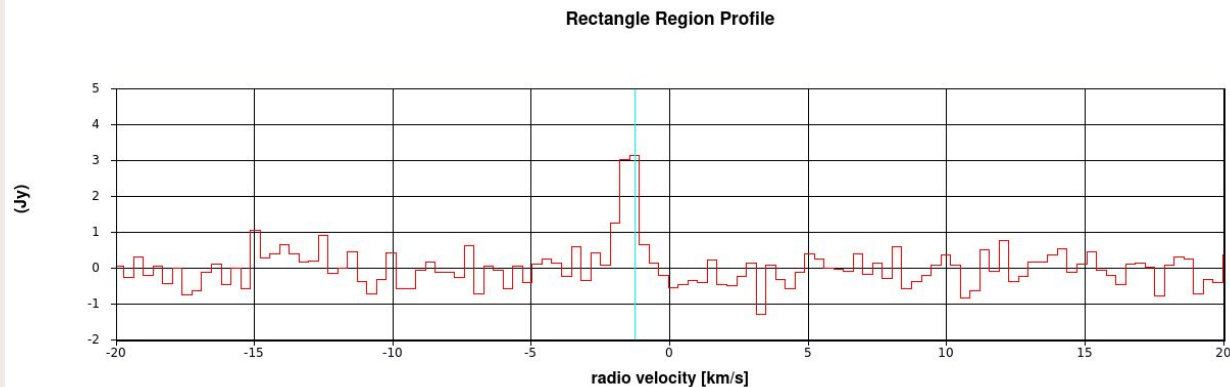
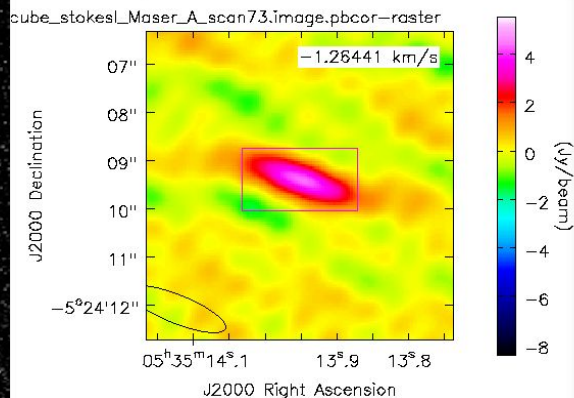
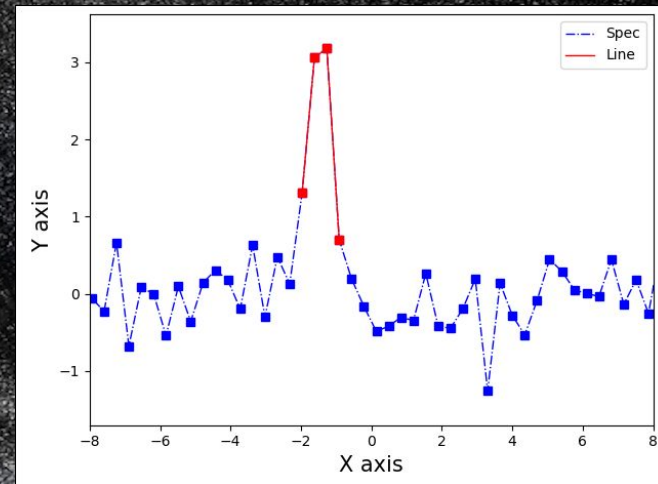
Epoch 3, Scan 73

- Lack of data
 - Not all telescopes took a reading
 - Left: Scan 73, Right: Scan 203



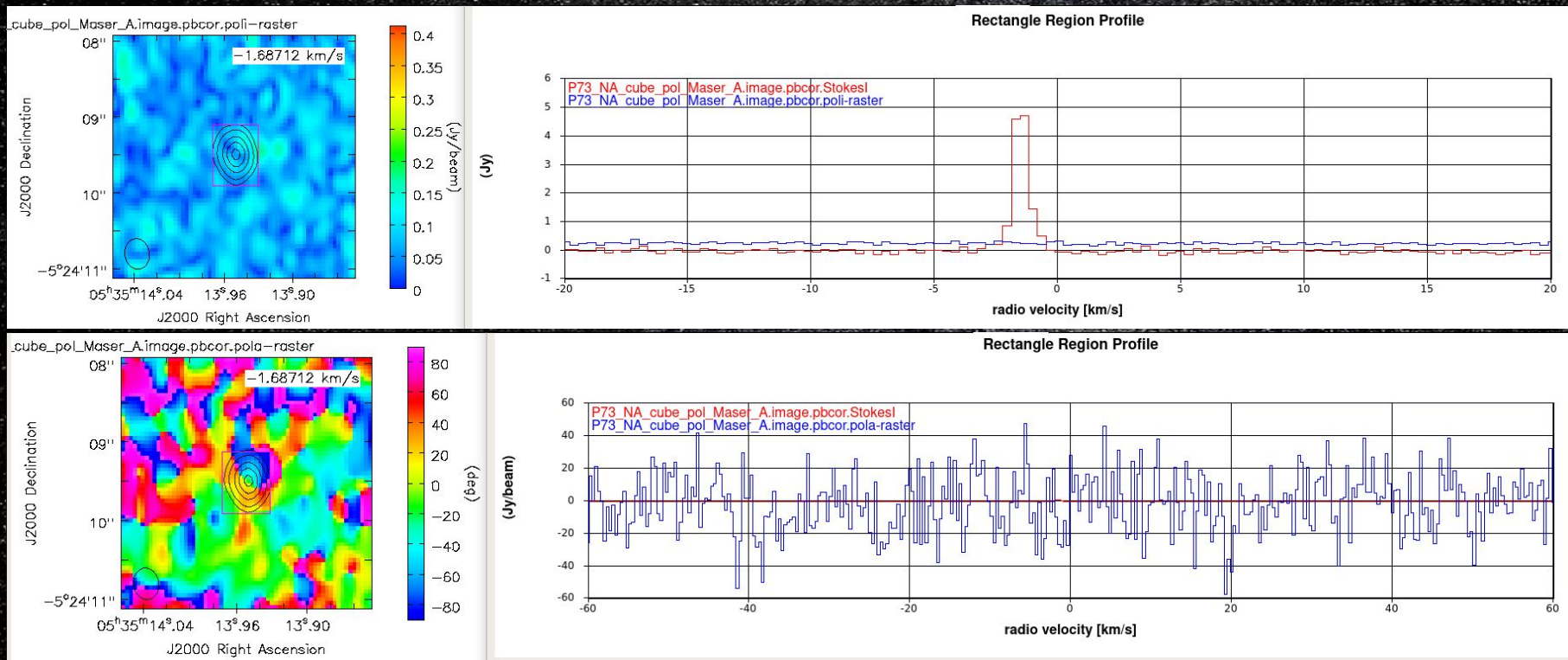
■ Scan 73

- More noise
- Less signal
- Lower signal-to-noise ratio



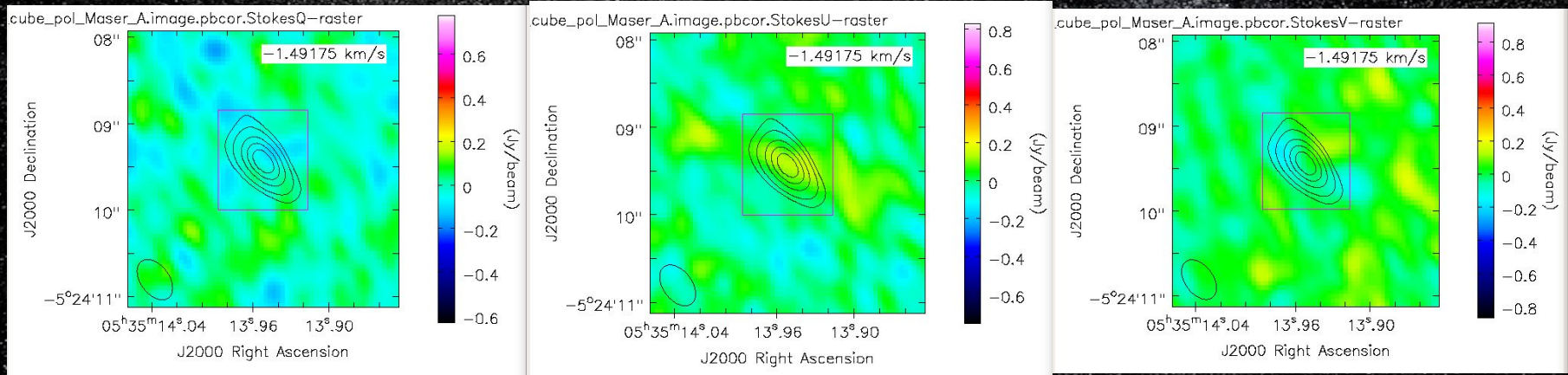
Objective 3

■ Lack of Polarization, Mix of Photonic Angles



Objective 3

- No dominant linear nor circular polarization found for any of the 20 epochs



Summary



- Constructed light-curve to investigate short-term maser variability
- General consistency for pointing averages and scans for most epochs
- Some evidence of fast variability, no evidence of polarization

Acknowledgements

- Thank you to:
 - My advisor Dr. Araya for mentoring me
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 - Dr. Dong, Dr. Trimm, and the rest of the SIR department

References

- Araya, E. D., Hofner, P., Goss, W. M., Kurtz, S., Richards, S., Linz, H., L. Olmi, & M. Sewilo. (2010). Quasi-periodic Formaldehyde Maser Flares in the Massive Protostellar Object IRAS 18566+0408. *The Astrophysical Journal Letters*, 717(2), L133–L137. <https://doi.org/10.1088/2041-8205/717/2/L133>
- Motte, F., Bontemps, S., & Louvet, F. (2018). High-Mass Star and Massive Cluster Formation in the Milky Way. *Annual Review of Astronomy and Astrophysics*, 56(1), 41–82. <https://doi.org/10.1146/annurev-astro-091916-055235>
- Sanchez-Tovar, E., Araya, E. D., Rosero, V., Hofner, P., & Kurtz, S. (2023). Broadband VLA Spectral-line Survey of a Sample of Ionized Jet Candidates. *The Astrophysical Journal Supplement Series*, 267(2), 43. <https://doi.org/10.3847/1538-4365/acdb5e>
- Sojka, G., VanZuiden, K., Araya, E., Busquet, G., Girart, J. M., Hofner, P., Fernández-López, M., Teixeira, P., Sánchez-Monge, A., Carrasco-González, C., Tobin, J., Rodríguez, T., & VOLS. (2024). Changes in the Spectral Zoo: Preliminary Results of Molecular Maser Variability in VOLS. *Bulletin of the AAS*, 56(7). <https://baas.aas.org/pub/2024n7i302p06/release/1>
- Voronkov, M. A., Sobolev, A. M., Ellingsen, S. P., & Ostrovskii, A. B. (2005). The 6.7- and 25-GHz methanol masers in OMC-1. *Monthly Notices of the Royal Astronomical Society*, 362(3), 995–1005. <https://doi.org/10.1111/j.1365-2966.2005.09374.x>